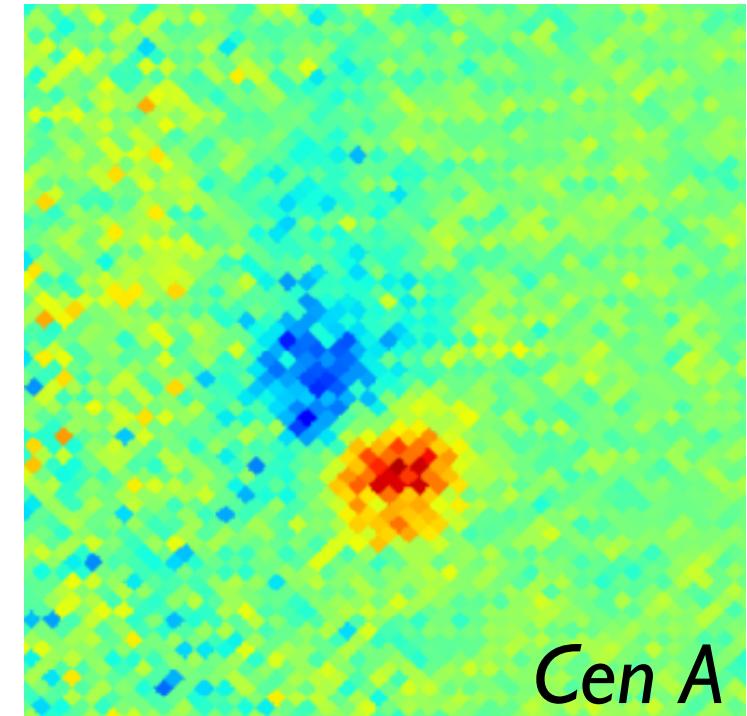
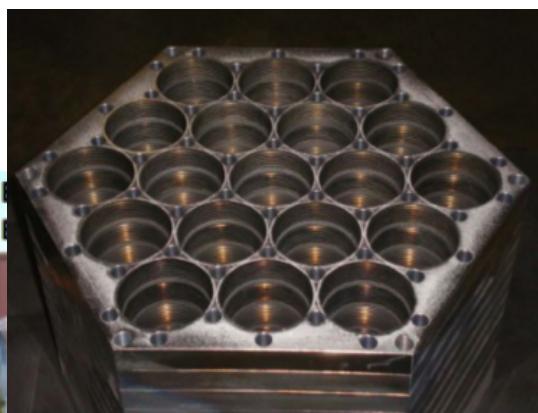
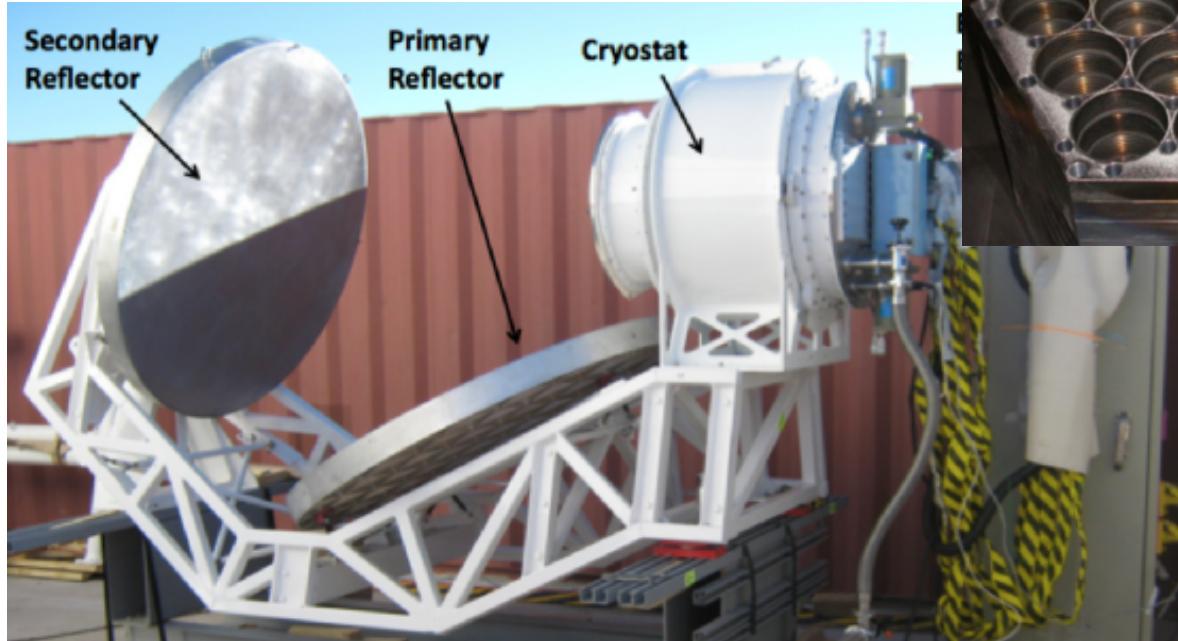


Polarization measurements of radio sources at 43 and 95 GHz with QUIET

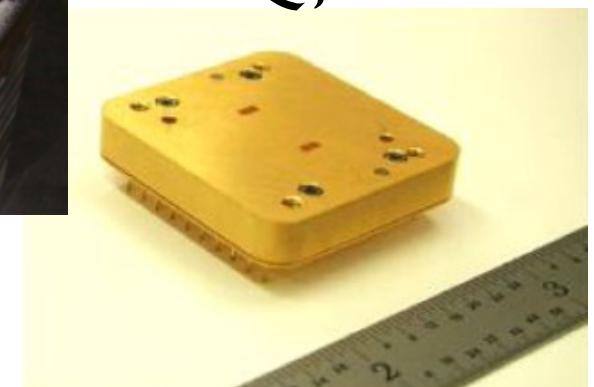
Kevin Huffenberger
Florida State University



QUIET = Q/U Imaging Experiment

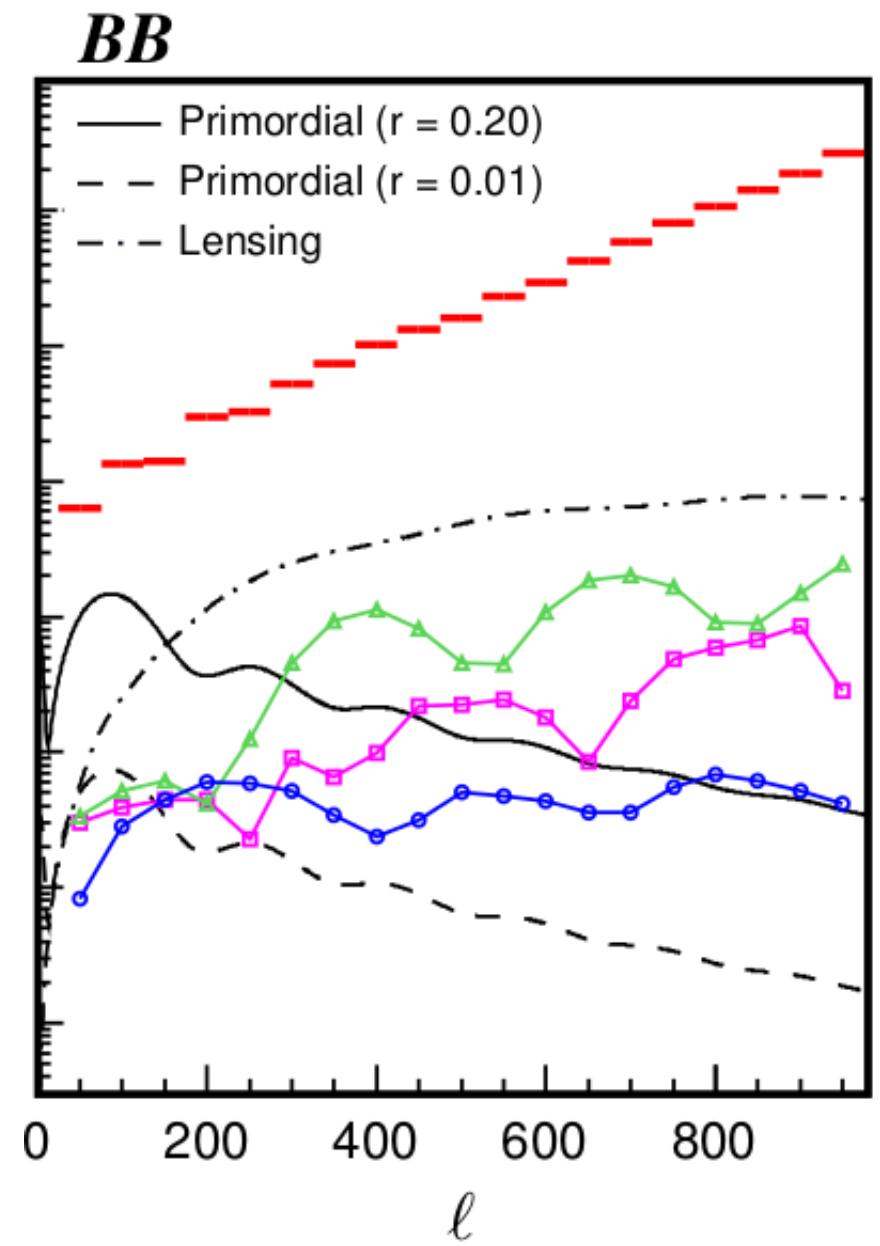
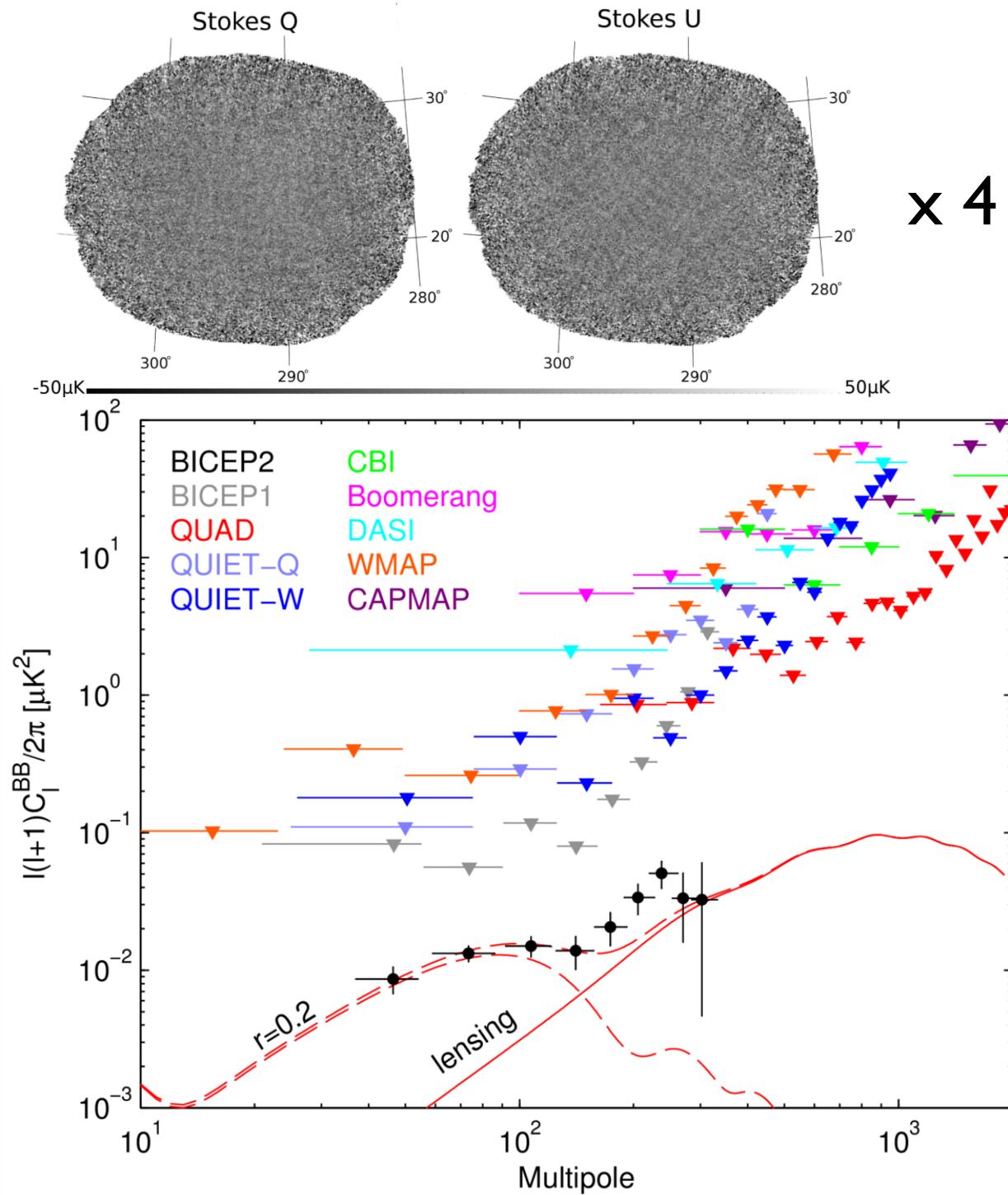


Q,U not I

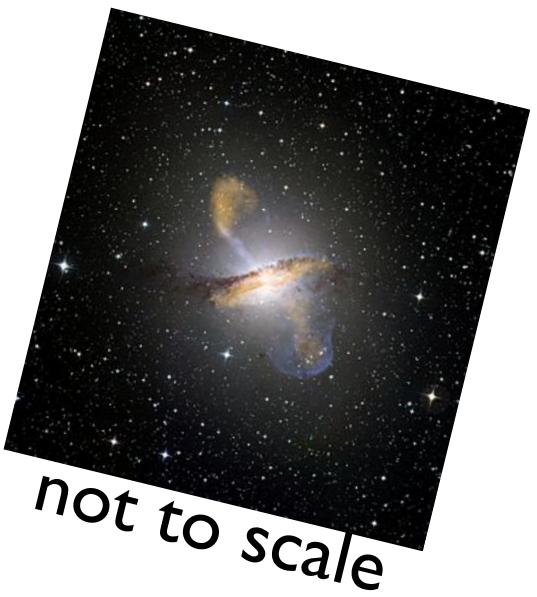


Q band: 43 GHz
W band: 95 GHz

QUIET CMB Results



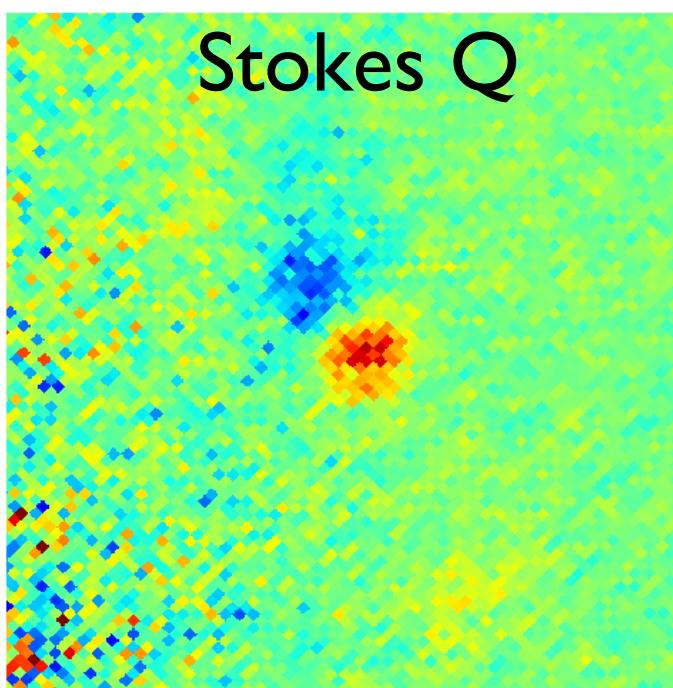
- Systematic Errors**
- Absolute Responsivity, Window Function
 - I to Q/U Leakage
 - Angle, Relative Responsivity, Pointing
 - Scan-synchronous Signal, Far Sidelobes



Cen A

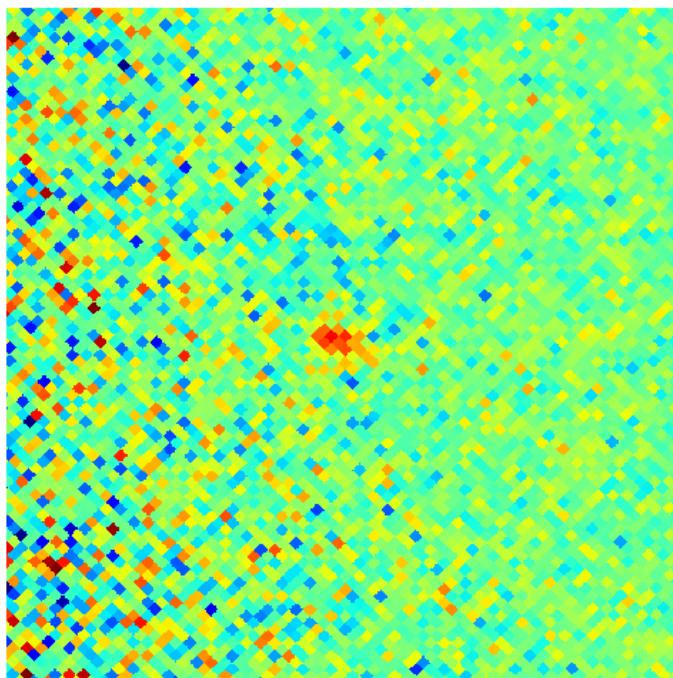
Q, U gal coords
CMB convention

Q band 43 GHz



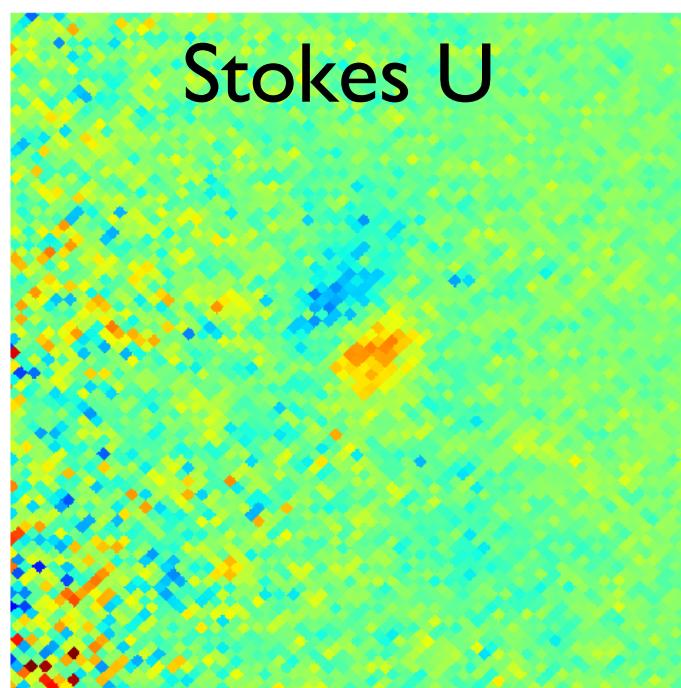
-1.8e+03 1.8e+03 μK

W band 95 GHz



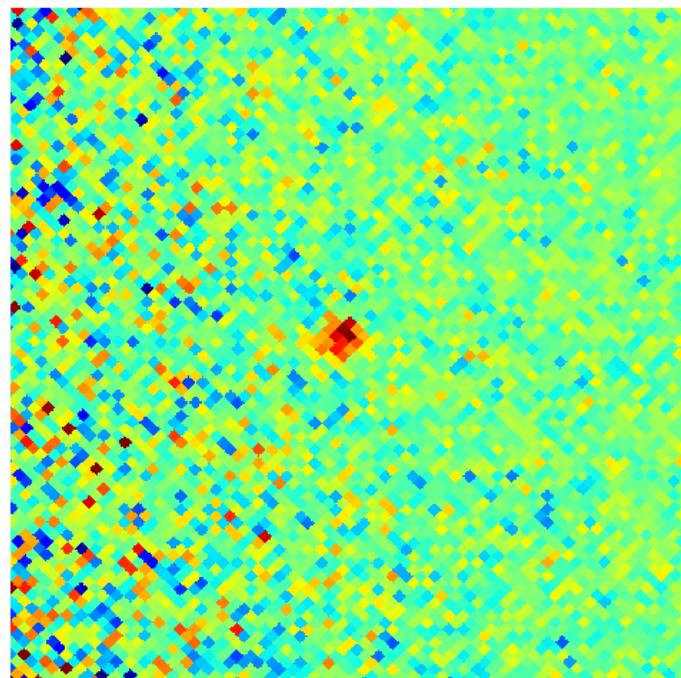
-500 500 μK

Stokes U



-1.8e+03 1.8e+03 μK

4°



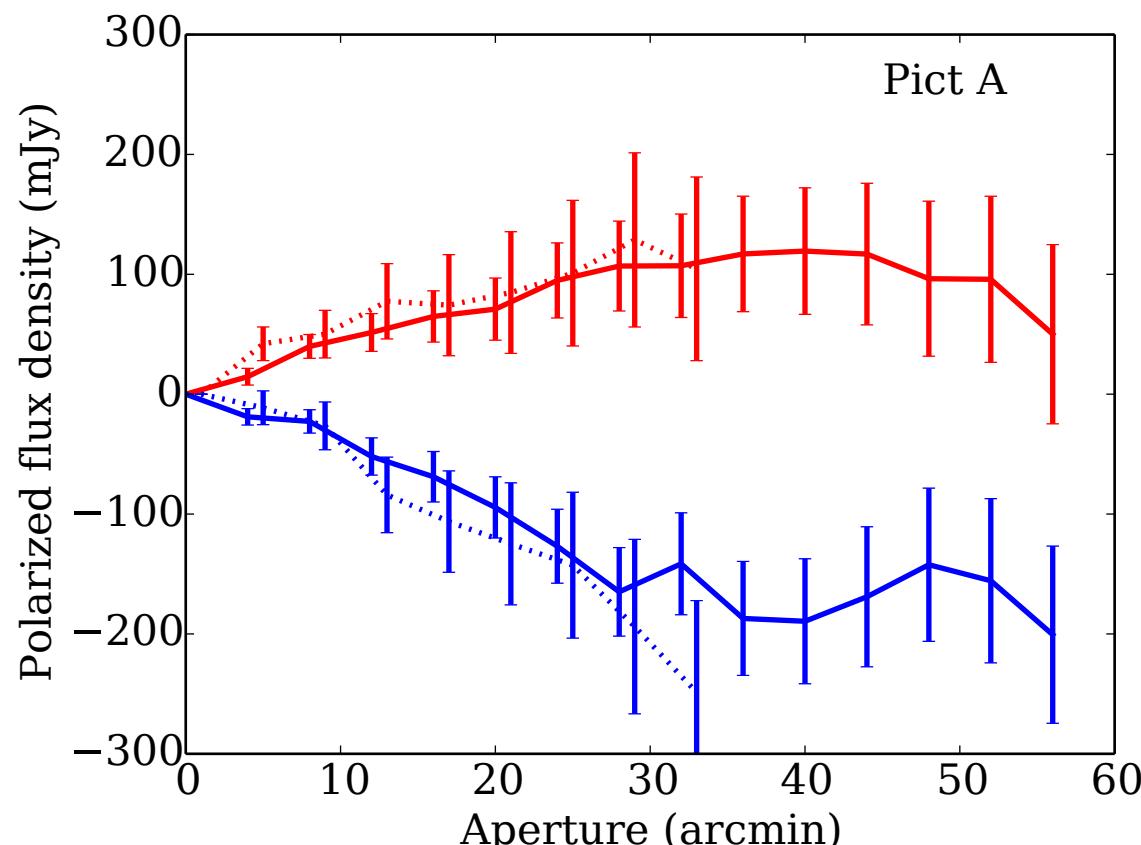
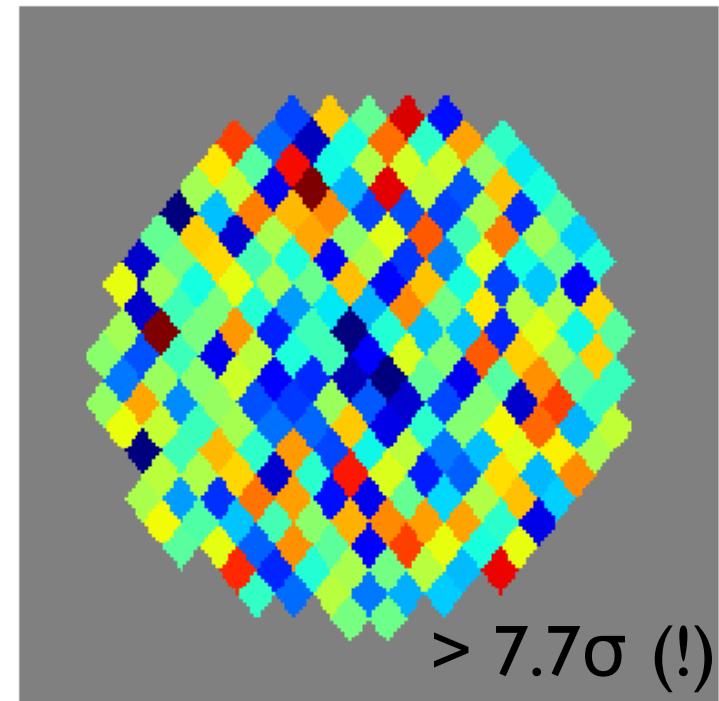
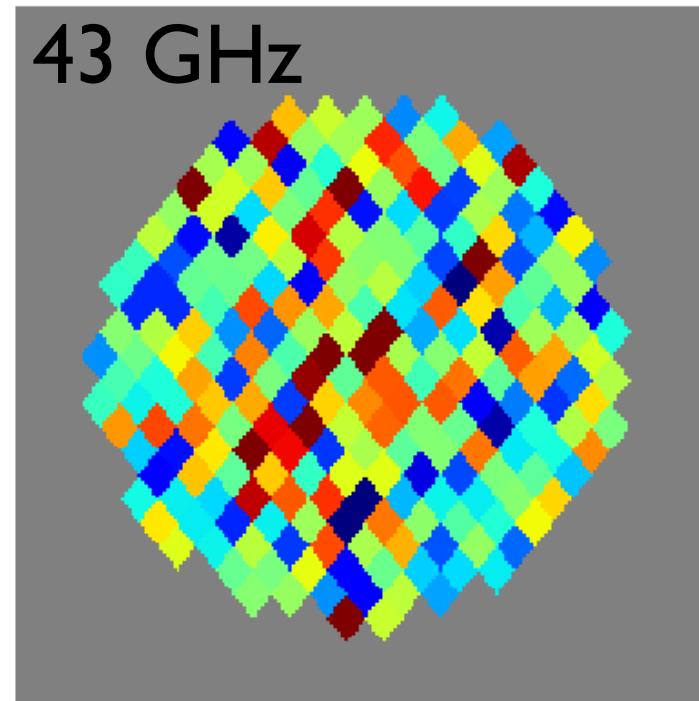
-500 500 μK

Pict A

Lobe sep.
 $\sim 8'$

1°

beams:
27' & 13'



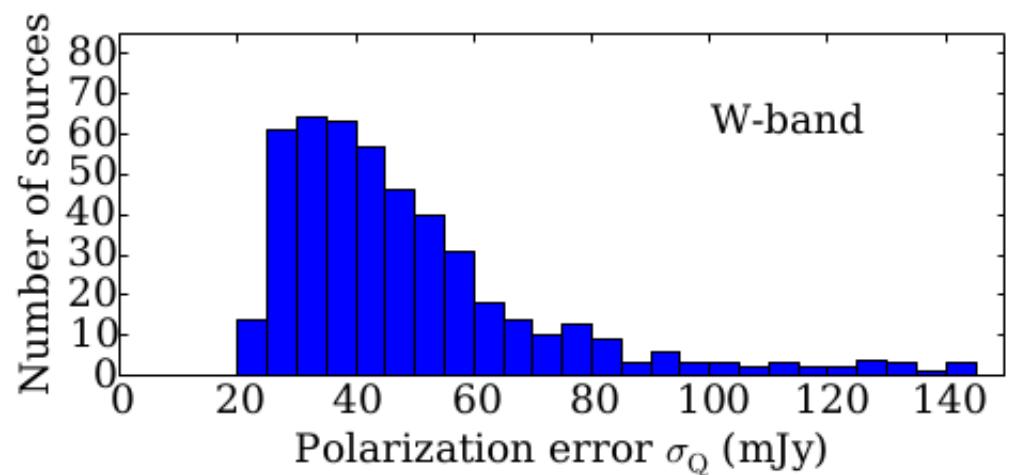
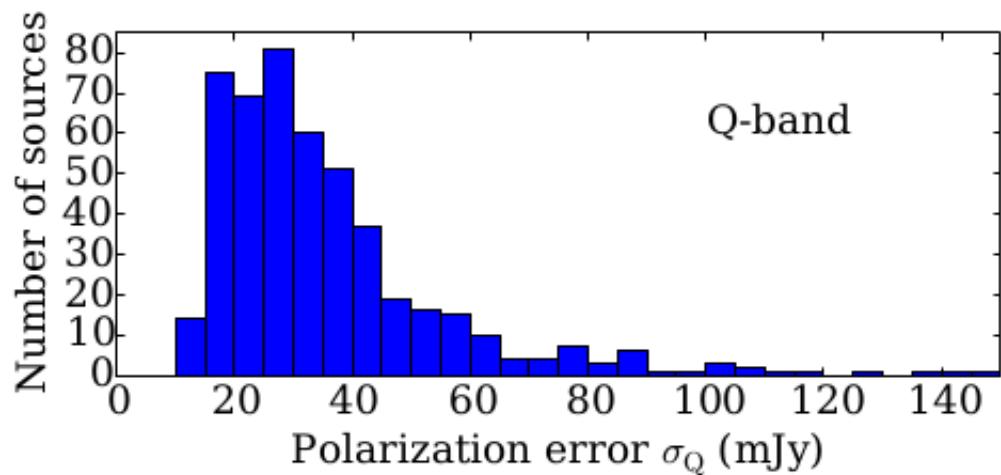
-50 μK 50

Stokes Q

Stokes U

Method for point-like sources

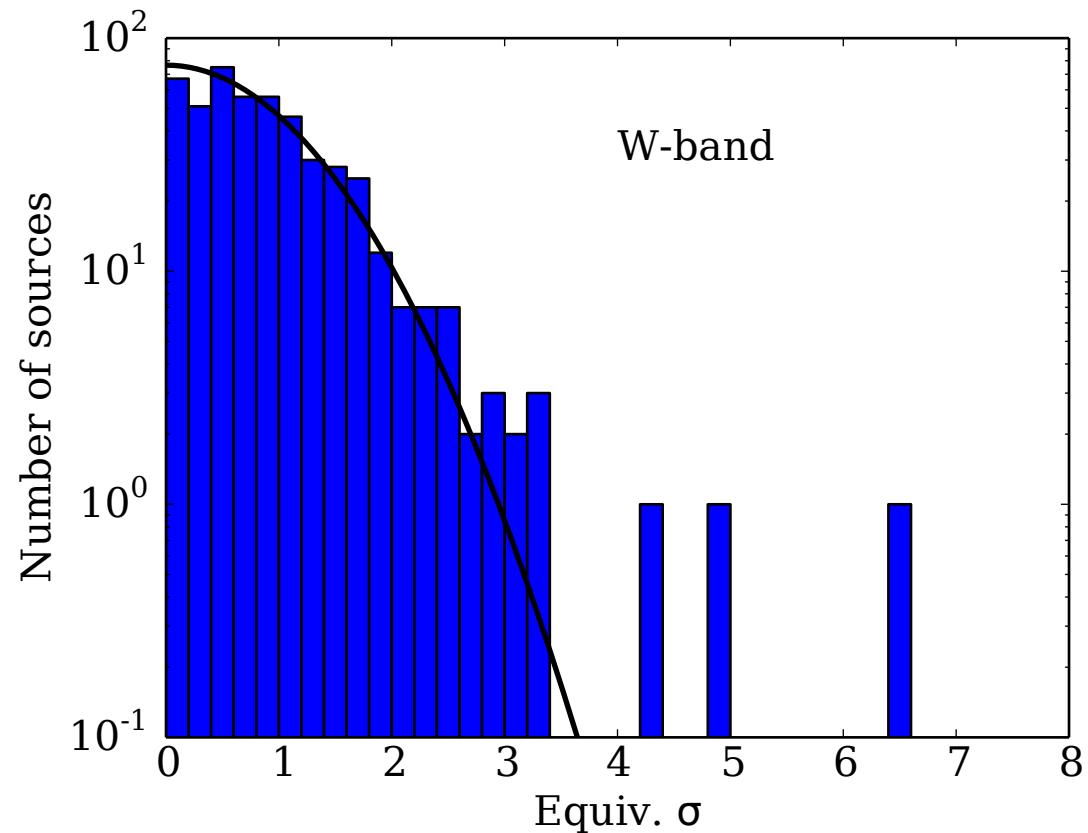
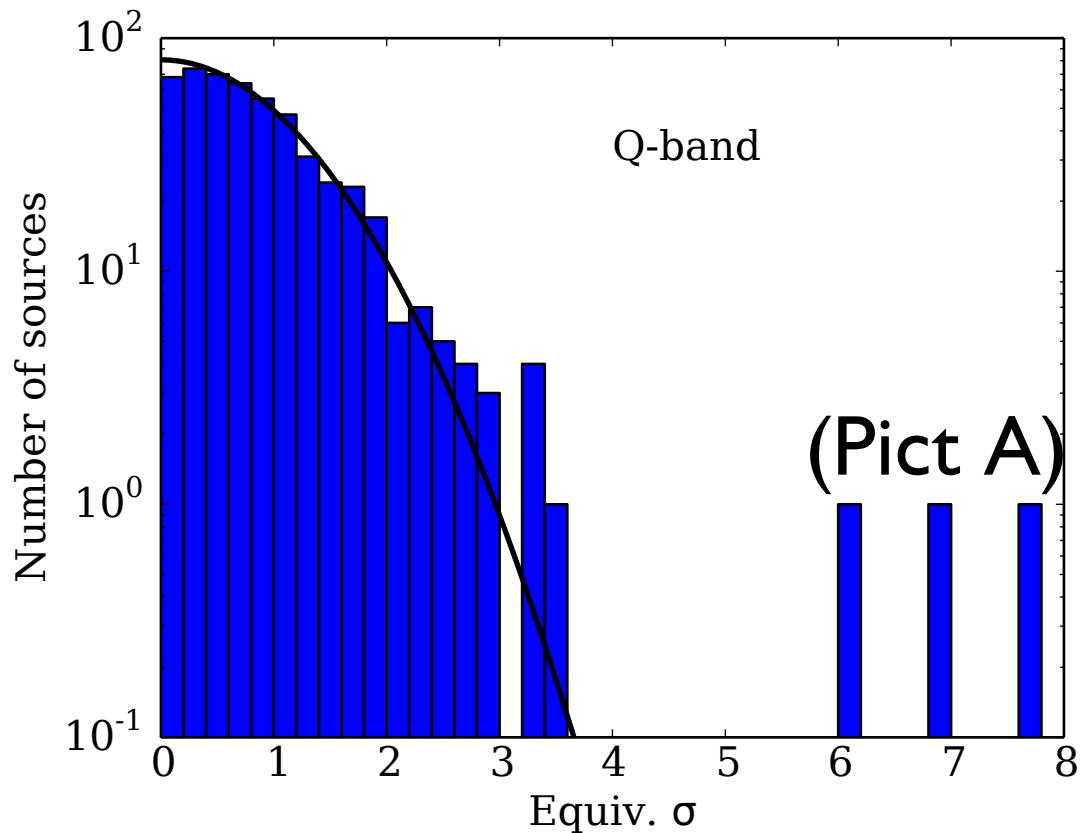
1. Build template for source shape (beam/psf + pix win.)
2. Measure Stokes Q,U on ~480 sources from Australia Telescope 20 GHz survey (all $> 5\sigma$) (Murphy+ 2010)



3. Evaluate probability by chance of measurement

4. MC, ML measurement of polarization magnitude

Number vs significance

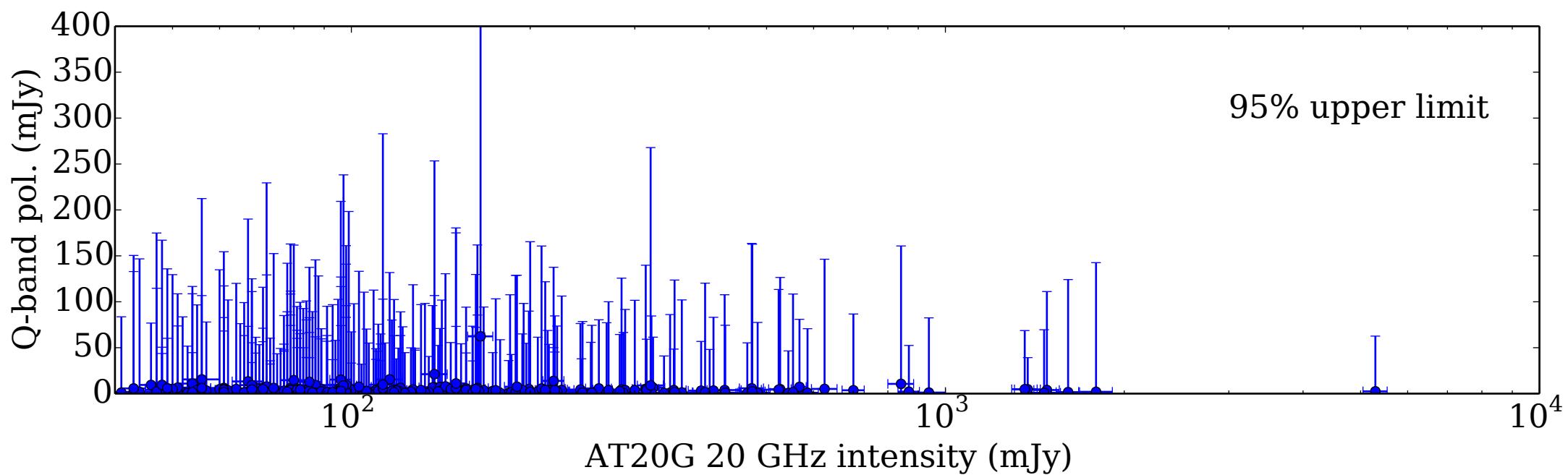
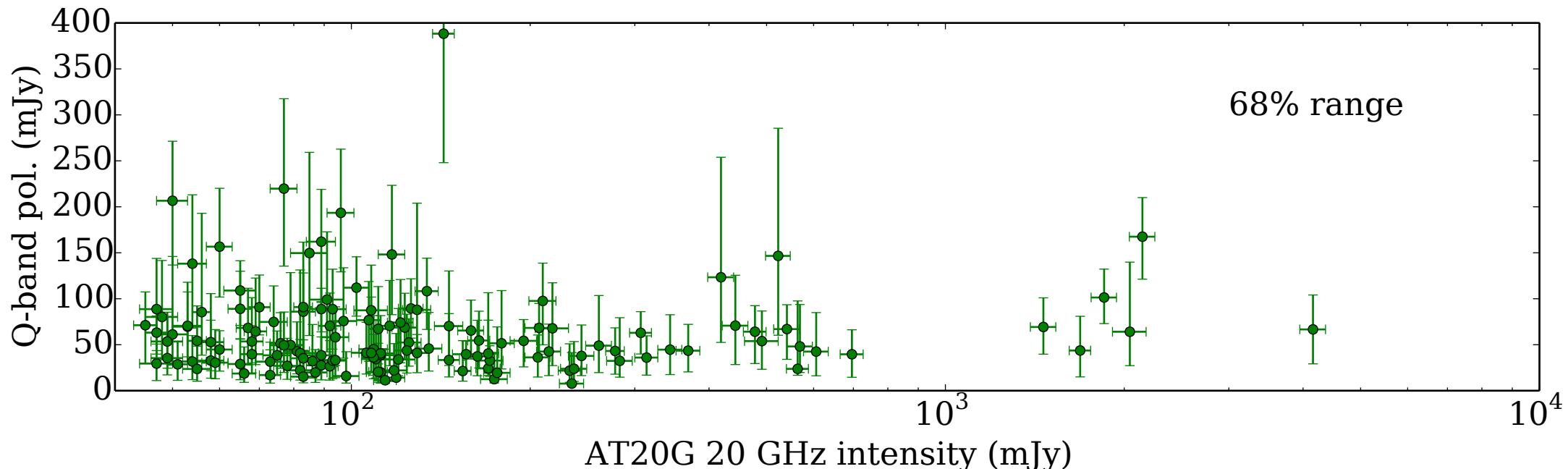


Expect 3.3 sources at $> 2.7 \sigma$ from noise alone

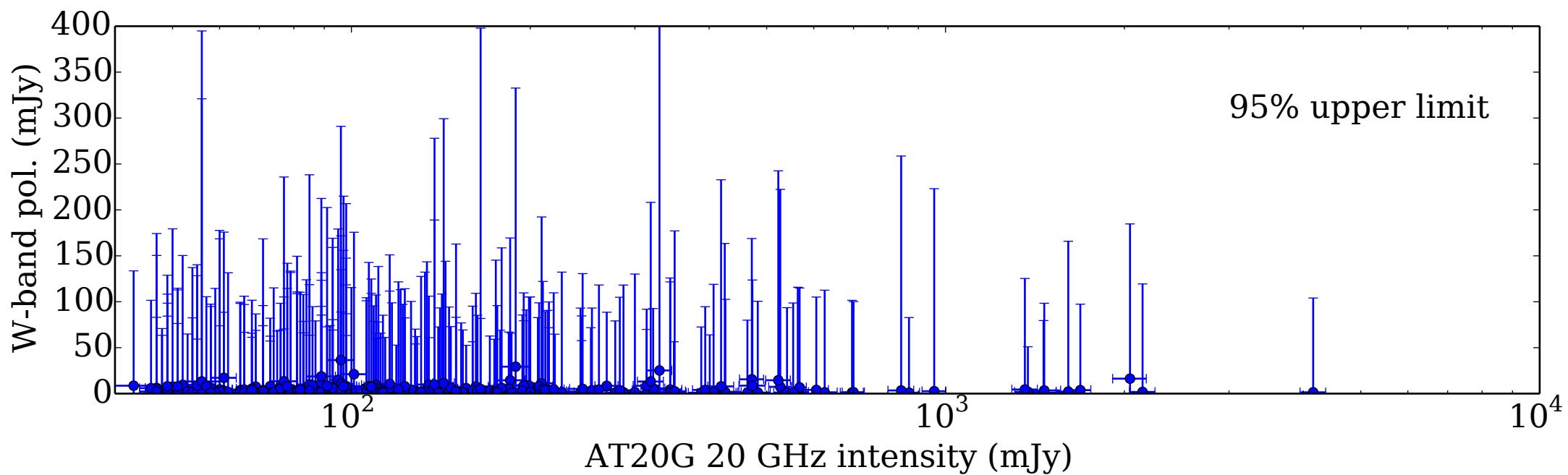
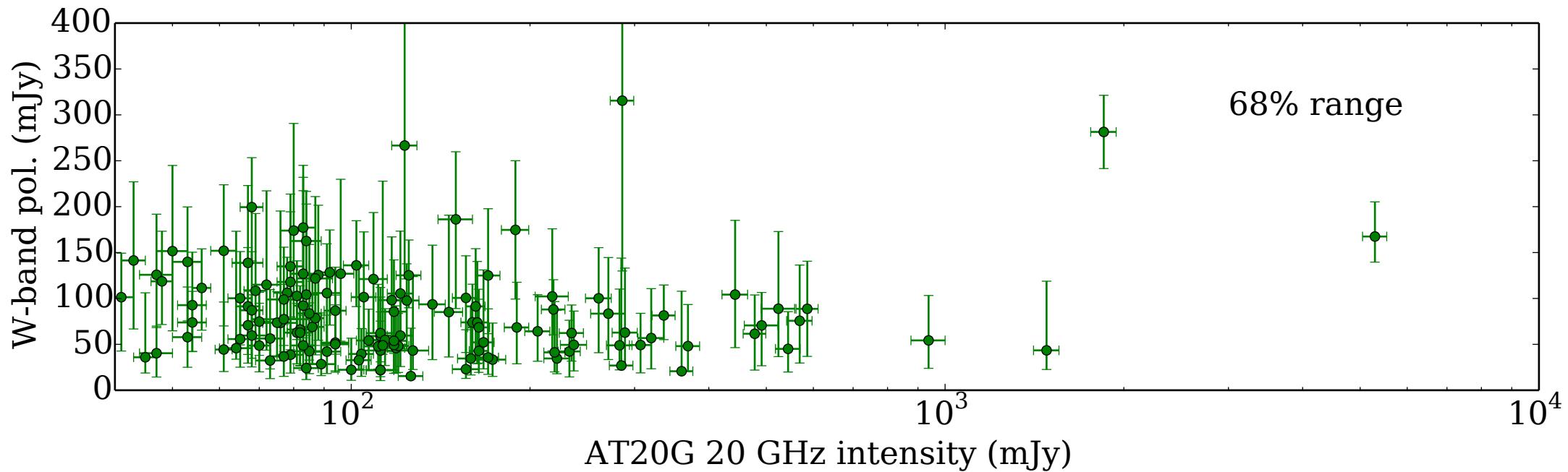
43 GHz: found 11

95 GHz: found 12

43 GHz polarization vs 20 GHz intensity



95 GHz polarization vs 20 GHz intensity



AT20GJ042840-375619, 3.22σ (Q-band), 6.58σ (W-band)

flux density (mJy)

3500

3000

2500

2000

1500

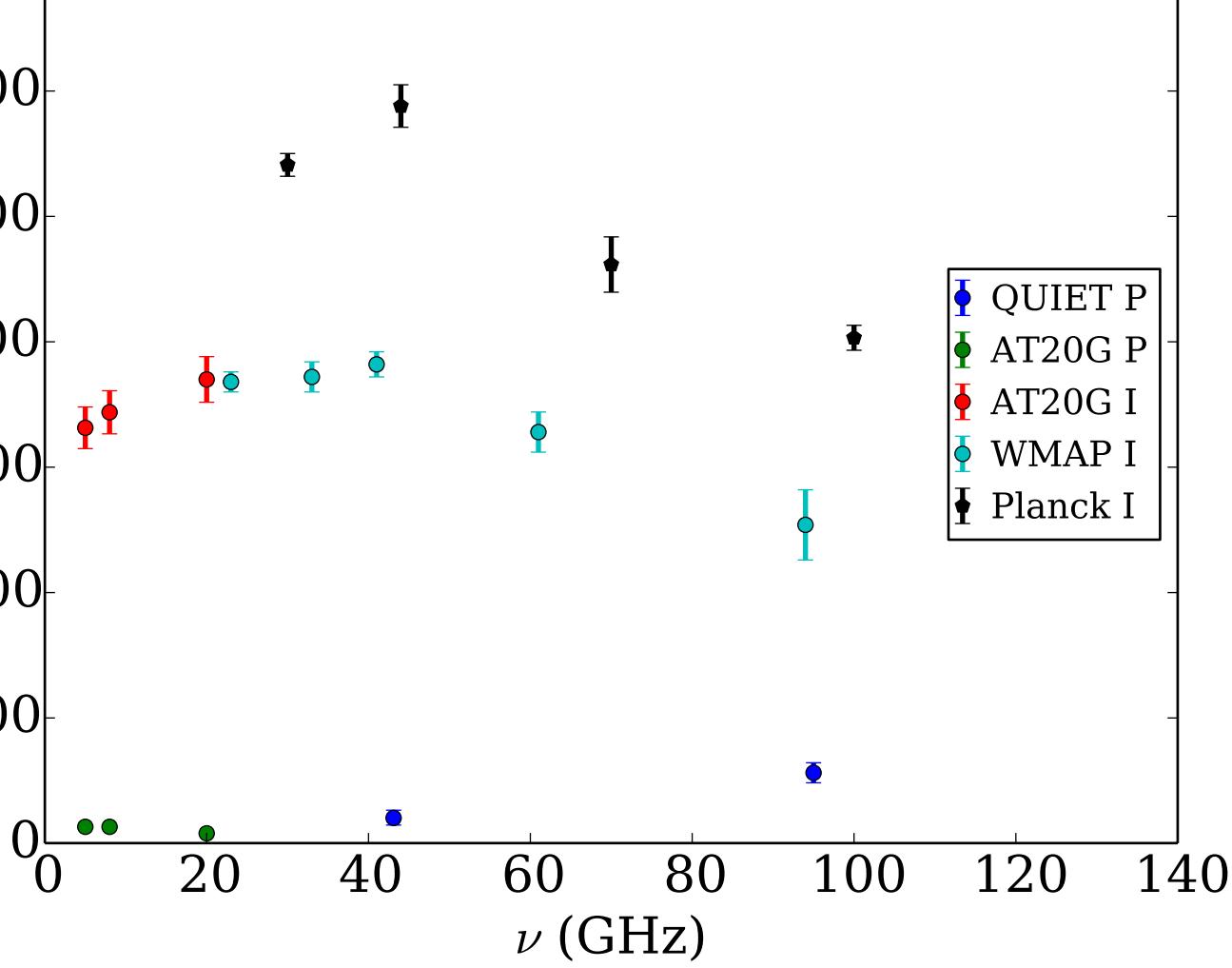
1000

500

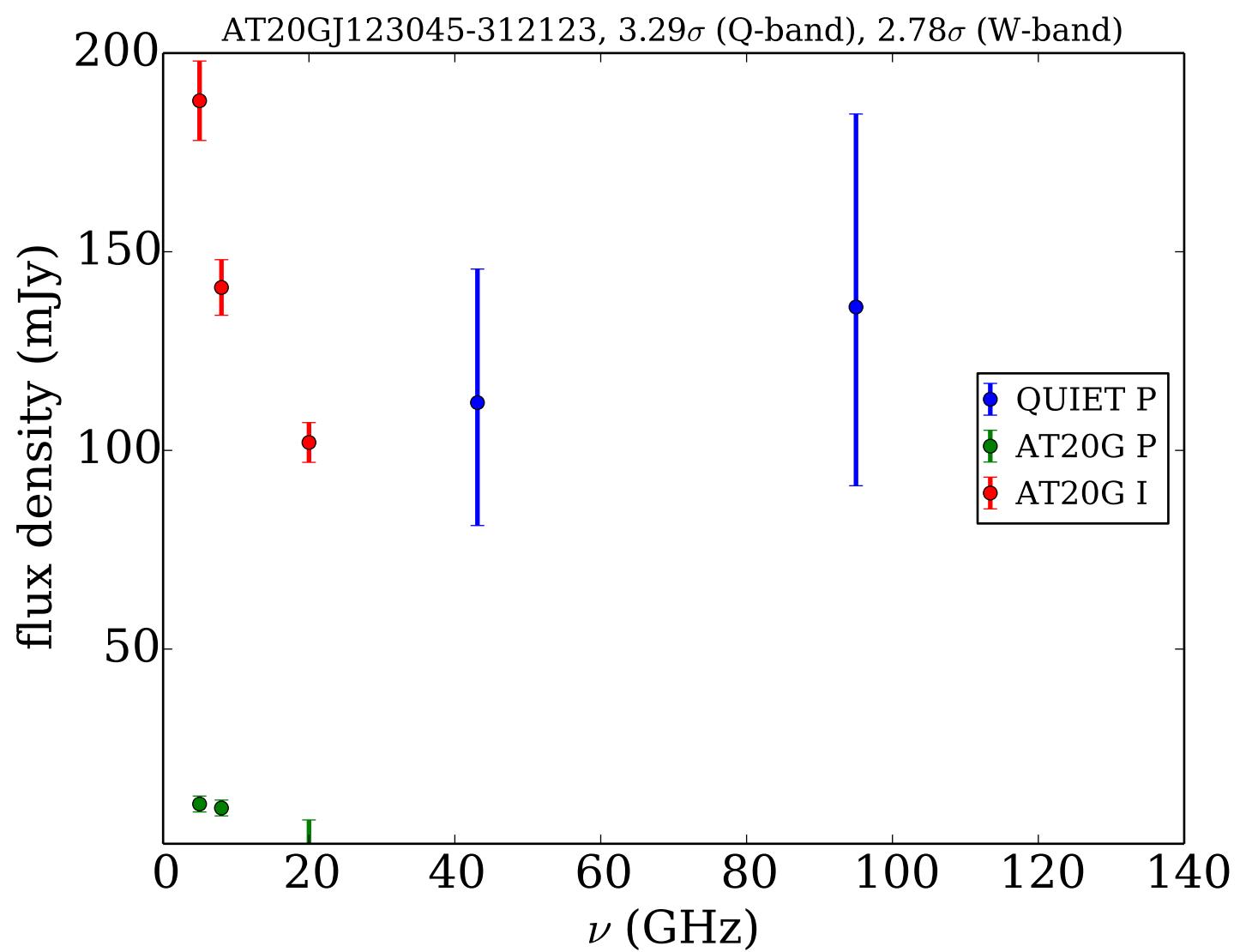
0

 ν (GHz)

- QUIET P
- AT20G P
- AT20G I
- WMAP I
- Planck I



AT20GJ123045-312123, 3.29σ (Q-band), 2.78σ (W-band)



Conclusions

QUIET measured polarization at the locations of ~480 radio sources at 43 GHz and 95 GHz.

Several detections of polarized emission; many upper limits.

No immediately clear trend between 20 GHz intensity and higher frequency polarization.

SEDs of sources are diverse and interesting.

Look for our catalog on arxiv this summer!